Letter

Magma–Shale Interaction in Large Igneous Provinces: Implications for Climate Warming and Sulfide Genesis

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Abstract

Large igneous provinces (LIPs) whose magma plumbing systems intersect sedimentary basins are linked to upheavals of Earth's carbon and sulfur cycles and thus climate and life history. However, the underlying mechanistic links between these phenomena are elusive. We address this knowledge gap through short time-scale petrological experiments (1200°C and 150 MPa) that explore interaction between basaltic melt and carbonaceous shale (mudstone) using starting materials from the Canadian High Arctic LIP and the Sverdrup Basin in which it intrudes. Here we show that entrainment of shale xenoliths in basaltic melt causes shale to shatter due to incipient thermal stress and devolatilization, which accelerates assimilation by increasing reactive surface area. Shale assimilation therefore facilitates transfer of sediment-derived volatile elements to LIP magma plumbing systems, whereupon carbon dominates the vapor phase while sulfur is partitioned into sulfide melt droplets. This study reveals that although carbon and sulfur are efficiently mobilized as a consequence of shale assimilation, sulfides can sequester sulfur—an important climate cooling agent—thus enhancing net emissions of climate warming greenhouse gases by shale-intersecting LIPs.

Keywords: magma-shale interaction, large igneous provinces, High Arctic LIP (HALIP), C-cycle perturbations, sulfide genesis

INTRODUCTION

Large igneous provinces (LIPs) are composed of massive emplacements of largely basaltic magma into Earth's lithosphere and onto the surface (Bryan & Ernst, 2008; Black et al., 2021). Many LIP events coincide with environmental change and mass extinctions (Courtillot & Renne, 2003; Bond & Grasby, 2017; Ernst & Youbi, 2017), but the mechanisms underpinning this correlation are not fully understood (Bond & Grasby, 2017; Black et al., 2021). The magnitude of carbon and sulfur released to the atmosphere is generally accepted as a key control on the severity of environmental impact of a LIP event (Jones et al., 2015). However, there is active debate concerning whether the release of primary, mantlederived volatiles by LIPs is sufficient to cause the global carbon cycle disruptions, climate change and biotic crises that mark mass extinctions or whether additional release of thermogenic, sediment-derived volatiles is required (e.g. Svensen et al., 2004, 2007, 2009; Ganino & Arndt, 2009; Yallup et al., 2013; Callegaro et al., 2021; Capriolo et al., 2021, 2022; Heimdal et al., 2021).

It is therefore necessary to clarify the mechanisms, dynamics and efficiency of volatile removal from sedimentary rocks that are

intersected by LIP plumbing systems. However, study of magmasediment interaction involving volatile-rich sedimentary rocks is fraught with obstacles such as poor outcrop preservation and accessibility. Sill-host rock contacts are often difficult to access and deeply weathered, sometimes necessitating collection of drill cores such as for the Siberian LIP (Callegaro et al., 2021) and the Brazilian portion of Central Atlantic Magmatic Province (Heimdal et al., 2019). Geochemical modelling of magma-sediment interaction is also limited, because while several platforms exist for modelling assimilation of silicate rocks, thermodynamic data for non-silicate materials such as carbonates, coal, organic matter, and evaporitic sulfates and halides are lacking (Heinonen et al., 2021). Efforts to study magma-sediment interaction through direct experimentation (using high pressure-temperature devices) has potential for generating unique mechanistic insights but has thus far largely been restricted to studies involving magmatic assimilation of carbonates (Freda et al., 1997, 2008, 2010; Iacono Marziano et al., 2007, 2008; Deegan et al., 2010, 2016; Mollo et al., 2010; Jolis et al., 2013; Carter & Dasgupta, 2015, 2016, 2018), although there is

growing interest in non-carbonate lithologies too (e.g. Iacono-Marziano *et al.*, 2017). Many of these previous studies utilized relatively long-duration experiments, with run times ranging from several hours to days, which yielded equilibrium melt-gas assemblages that can be difficult to interpret in terms of process.

Short-duration (seconds to minutes) high pressure-temperature (i.e. 'kinetic') experiments that better replicate the onset of magma-sediment interaction emerged in the early 2010s (Deegan et al., 2010) and paved the way for close examination of the rapid transformations that occur when basaltic melt first comes into contact with sedimentary rock. In this paper, we apply a similar philosophy to simulate entrainment of carbonaceous mudstone (hereafter referred to as 'shale') xenoliths in basaltic melt and explore incipient magma-shale interaction at LIPs. The experiments preserve textural and chemical information that is largely obliterated in longer duration experiments as well as in many natural contexts, which rarely freeze-in prograde reaction paths (Fig. 1a). Magma-shale interaction moreover involves a sulfur component, which extends its relevance to sulfide ore formation. It is generally accepted that sulfide formation in LIP-hosted Ni–Cu–platinum group element (PGE) deposits often involved incorporation of sedimentary crustal sulfur during magma emplacement (Lesher, 2019). Since the solubility of sulfide in basaltic melts is typically low, most of the sulfur needed to constitute Ni-Cu-PGE ore bodies calls for addition of crustal sulfur as sulfide 'xenomelt(s)', i.e. a foreign melt derived from crustal rocks or xenoliths (Lesher, 2017, 2019). The crustal origin of sulfur is supported by isotopic and geochemical data, but uncertainty remains surrounding how sulfide xenomelts form and are transported in ore-forming systems (Hayes et al., 2015; Lesher, 2017, 2019). These are questions that short-duration interaction experiments can also potentially clarify.

FIELD AREA, STARTING MATERIALS AND METHODS

Here we employ the Cretaceous High Arctic LIP (HALIP) as a case study because its mafic sill province, representing the frozen remains of the conduit system formerly supplying flood basalts to the surface, is spectacularly preserved and exposed in the Canadian Arctic (Jowitt et al., 2014; Evenchick et al., 2015; Deegan et al., 2018; Bédard et al., 2021a, 2021b). Much of the Canadian portion of the HALIP was emplaced at ca. 120 Ma into the Sverdrup Basin, a sedimentary depocenter filled with up to 13 km of siliciclastic, evaporitic and carbonaceous strata of Carboniferous to Paleogene age (Embry & Beauchamp, 2019). HALIP sills that invaded Sverdrup Basin sedimentary rocks collectively influenced basin-scale thermal evolution and development of regional oil and gas resources (Jones et al., 2007; Goodarzi et al., 2019) and possibly also sulfide ore formation (Jowitt et al., 2014; Saumur et al., 2016). The Middle Triassic Murray Harbour Formation forms part of the Sverdrup Basin fill and is dominated by black shale and siltstone, with subordinate calcareous and phosphatic interbeds. This unit is rich in Type II kerogen and is considered to be the primary source rock for hydrocarbon discoveries in the Sverdrup Basin (Brooks et al., 1992; Kondla et al., 2015).

We utilized natural HALIP magmatic rock and Murray Harbour Formation shale as the experimental starting materials. The experimental conditions and compositions of the starting materials are provided in the **Supporting Data File**. Details of the experimental method and analytical techniques applied to the experimental products are provided in full in the **Supporting Information**. To summarize, the magmatic starting material is a pristine, non-cumulate mafic rock sample from a ca. 30 m thick sill with 6.6 wt % MgO and 0.15 wt % S. The shale is a finely clastic mudstone collected from a site over 60 m perpendicular distance away from any observable sill contacts and contains 4.7 wt % total organic carbon and 0.5 wt % total S. The shale is principally composed of quartz (56 wt %), calcite (24 wt %), biotite (9 wt %), dolomite (4 wt %) and pyrite (1 wt %) and has a loss on ignition value of 21.6 wt %, indicating that roughly one fifth of its mass comprises volatile compounds including sulfur from pyrite and organic matter plus inorganic and organic carbon.

The magmatic rock was converted to glass, both an anhydrous and a mildly hydrated aliquot, pulverized in an agate mill and loaded into platinum capsules (ca. 25 to 35 mg of powdered glass per experiment) along with a solid fragment of shale (ca. 4 to 5 mg per experiment). Each capsule was positioned in a 19-25 mm NaCl-crushable MgO-borosilicate glass assembly employing a graphite furnace and was then pressurized to 150 MPa in a low-P calibrated non-end loaded piston cylinder device (see Masotta et al., 2012). The assemblies were then heated to 1200°C at a rate of 100°C per minute and held at the set point temperature for durations of 0 s, 300 s and 600 s while being continuously monitored for any P-T fluctuations. The experimental P-T conditions are considered appropriate for injection of HALIP tholeiites at or slightly above their liquidus (1100 to 1200°C) at ca. 5 km depth in the Sverdrup Basin (see Bédard et al., 2021a). The experiments were terminated by shutting down the power source whereafter they were isobarically quenched at a rate of ca. 100°C per second in the first 5 seconds (to the glass transition) and ca. 33°C per second thereafter. Fast, isobaric quenching is important to preserve textures and glass (≈melt) that formed in situ within the time window of xenolith dissolution and melting (Fig. 1a). After quenching, the capsules were retrieved, cast in epoxy, polished and inspected using secondary electron microscopy (SEM). Quantitative analysis of major element oxides and sulfur was subsequently performed utilizing electron probe microanalysis (EPMA) while volatile species were determined using confocal Raman microspectroscopy (see Supporting Information for analytical details).

EXPERIMENTAL RESULTS

An extended description of the experimental products is provided in the **Supporting Information**, and the compositions of the glasses and sulfides in the experimental products are provided in the **Supporting Data File**. An overview of the textures of the experimental products is presented in Fig. 1 as a series of SEM mosaic images and a summary of the compositions of the experimental glasses, sulfides and vapor bubbles are provided in Figs 2 and 3. In brief, the experiments document features of incipient magmashale interaction, including the following processes:

- (i) Dissolution of shale into the host melt, manifest as disequilibrium textures and crustally contaminated, 'modified' melts (Figs 1 and 2).
- (ii) Formation of modified melts that are initially restricted to patches within the shale fragment but eventually form a boundary layer surrounding shale. Calcium is enriched (relative to the starting HALIP basalt) in modified glass in all experiments while silica, potassium and sulfur are often enriched too (Fig. 2).
- (iii) Shattering and degassing of shale and formation of abundant vapor bubbles containing C (disordered graphite), CO and CO₂ (Fig. 3).







Fig. 1. (a) Relative timescales for melting host rocks and xenoliths by magmatic intrusions (modified after Robertson et al., 2015). This study targets the timeframe of cm- to m-scale xenolith melting in order to capture evidence of interaction processes that are only rarely preserved in nature. (b) to (g) Backscattered electron (BSE) mosaic images of experimental run products. Intense degassing is evidenced by abundant small vesicles that permeate the run products (especially in the 0 s runs). Sulfides are generally found at magma–shale interfaces. Runs held for 600 s show a more advanced stage of shale dissolution into the host melt. Scale bars are 1 mm.

(iv) Formation of sulfide mineralizations (≈sulfide xenomelts) at magma–shale interfaces (Figs 2 and 3).

The experiments therefore reveal that contaminated melts and carbon volatiles are generated rapidly as a consequence of shale assimilation but that sulfides effectively sequester sulfur, a climate cooling agent, thus enhancing net emissions of greenhouse gases by LIPs that intersect shale-bearing sedimentary basins (as discussed below).

DISCUSSION

Magma-sediment interaction in microcosm

Our experiments have replicated magma–shale interaction in microcosm, i.e. the experiments encapsulate in miniature the characteristics of a much larger system. Earlier, similarly designed experiments showed that magma–carbonate interaction results in a low-viscosity, Ca-rich compositional boundary layer at the reaction site and a voluminous C–O–H (carbon–oxygen–hydrogen)



Fig. 2. (a) Sketch of experiment SH-7G. To the right are EPMA glass data for SiO₂, CaO, K₂O and S, respectively. **(b)** Sketch of experiment SH-20G-H. To the right are EPMA SiO₂ and CaO maps. Small black squares on the maps represent EPMA glass analysis sites and numeric values are in weight %. **(c)** EPMA Fe, S, Cu and Ni maps for sulfides at the melt-shale interface (white box on SiO₂ map). Numeric values are in weight %. Abbreviations: Qz, quartz; Rxn, reaction; SCSS, sulfur concentration at sulfide saturation; SM, starting material. All EPMA glass data are normalized to 100% and the glass and sulfide data are reported in the Supporting Data File.

vapor phase (e.g. Deegan et al., 2010). We find evidence of this process in our experiments, but shale assimilation is complicated by the fact that shale contains silicate minerals and organic matter in addition to carbonate. Previous experimental work involving siliceous shale and/or metapelite showed that siliceous partial melts formed at ca. 700 to 800°C at pressure equivalent to ca. 200 MPa (Wyllie & Tuttle, 1961; Erdmann et al., 2007). Recent shale heating experiments at 200 MPa have also shown that



Fig. 3. (a) BSE mosaic image of experiment SH-17G-H showing the locations of Raman analysis sites. **(b)** BSE image of the shale-melt interface showing the texture of partly dissolved shale, sulfides along the shale-melt boundary and abundant bubbles in the glass. This is a surface image so the bubbles that are visible were opened during sample polishing. Raman analyses were conducted on unopened bubbles beneath the sample surface. **(c)** Raman spectra for the analysis sites are shown to the right, with C, CO, CO₂, O₂ and N₂ bands identified (note that O₂ and N₂ are likely derived from air introduced during capsule loading, but O₂ may also derive from decomposition of oxidized carbon species by laser heating; see Supporting Information for further details).

C–O–H–S fluids are generated at temperatures ≤700°C and that silicate melt is produced from ca. 800°C (Virtanen et al., 2021). None of these previous studies simulated interaction between shale and basaltic melt, but there are nevertheless some broad similarities to our work. For instance, two of our experiments contain a separated high-silica glass phase (Fig. 1b,c), which we suggest is the product of quartz breakdown at <1200°C during experiment heat-up. These segregated high-silica glass domains are not present in the mildly hydrated experiments, which suggest that melt mixing was more efficient at low temperatures (during heat-up) when there was water added to the basaltic melt.

Importantly, we observed a portion of Ca-enriched, compositionally modified (crustally contaminated) glass near the reaction site in all of our magma–shale interaction experiments (Fig. 2a,b). The at times extremely high CaO content of this glass (up to 45 wt %) indicates that the carbonate components of the shale dissolved rapidly into the host melt. In many experiments, the compositional boundary layers are also enriched in silica, potassium and sulfur from dissolution of quartz, micas and pyrite (Fig. 2a,b). These silica-rich contaminated melts are particularly significant because in natural systems they would potentially crystallize zircon, which would enable U–Pb geochronology of basaltic magmas injected into shale (Gaynor et al., 2022).

In addition to demonstrating local contamination phenomena, the experiments provide compelling evidence for transfer of carbon volatiles from shale to melt. Evidence of a vapor phase is preserved in the experiments as bubbles that coalesce and migrate within minutes. Our limited number of experiments does not allow us to quantify how vesicle volume changes as a function of reaction time, but the shale appears to undergo a significant volume change, with increasingly smaller fragments visible over time (Fig. 1). The volatile mix generated is comprised dominantly of C (disordered graphite), CO and CO_2 as revealed by spectroscopic analysis of small (ca. 20 μ m diameter) subsurface bubbles in a zero-time experiment (Fig. 3). Methane and water may be expected too (cf. Capriolo et al., 2021), but further analyses would be required to verify their presence. Sulfur likewise cannot be ruled out as a vapor phase although it appears to be largely sequestered into sulfides that dot the melt-shale interface.

Thermal stresses and concentration of volatiles produced along cleavage planes in shale would trigger a localized volume change, which could lead to fragmentation of shale and open up new surface area for reaction, triggering a short-lived 'runaway' effect with respect to volatile expulsion. Since the host melt is virtually incapable of storing dissolved CO₂ at 150 MPa (the solubility of CO₂ in basaltic magma at 150 MPa is ca. 0.07 wt %; Newman & Lowenstern, 2002), the system would quickly become fluid oversaturated and most of the carbon in the system would be expelled as a free vapor or fluid phase capable of hydrofracturing roof and wall rocks. Notably, thermodynamic models predict that assimilation of as little as 0.6 wt % organic matter will cause a doubling of the total volatile load of a magma and produce CO-dominated gases (Iacono-Marziano et al., 2012). We have now verified these predictions via direct analysis of bubbles (Fig. 3) and provide unique empirical evidence for CO gas formation during magma-shale interaction.

Ground truthing the experimental results

There are several natural examples that mirror our experimental results, which we briefly summarize here: (1) shale xenoliths reported from Nuussuaq (Greenland) and the Duluth Complex (USA) show evidence for extensive reaction and element exchange with their host melt (Pedersen & Larsen, 2006; Samalens et al., 2017); (2) studies of mafic sills injected into shale in Skye (Scotland) show that shale close to sill contacts displays abundant vesicles, partial melt textures with glassy streaks enriched in sulfur, silica and potassium, Fe-S mineralizations and evidence for sulfur loss (Lindgren & Parnell, 2006; Yallup et al., 2013); and (3) sill-shale contacts in the Canadian HALIP possess sulfides with textures similar to those reported here (see Supporting Information). These natural cases underscore both the utility of our experimental approach in understanding contact phenomena and the dynamic, rapidly evolving nature of magma-sediment interaction.



Fig. 4. (a) Illustration of a LIP plumbing system intersecting a sedimentary basin. Sill injection into shale generates sediment-derived volatiles (fluids and/or vapor). If these volatiles enter the atmosphere and/or ocean through a vent complex or eruption, they can trigger climate warming and ocean anoxia. **(b)** Generalized model of sill-host rock interaction with a close-up of incipient magma-shale interaction as revealed by experiments. Magma-shale interaction causes formation of carbon volatiles, crustally contaminated melts and sulfide xenomelts. Note that contact aureoles reach a maximum temperature (max. aureole T) of ~650 °C depending on multiple factors including sill temperature, ambient temperature, sill thickness and host rock type and are a source of volatiles not directly studied in our experiments. Sill liquidus and solidus curves are estimated based on HALIP petrological models (Bédard et al., 2021a). Abbreviations: devol., devolatilization; Liq, liquidus; Sol, solidus; S. Basin, sedimentary basin; T, temperature.

WIDER IMPLICATIONS Formation of sulfide xenomelts

Thermodynamic models have shown that as little as 0.1 wt % magmatic assimilation of organic matter (CH) will cause oxygen fugacity (fO₂) to decrease by more than two log-units (Iacono-Marziano et al., 2012). Magmatic assimilation of shale therefore facilitates a reducing melt environment while also transferring sulfur from shale to the host melt. There may, however, be local and/or transient variations in fO₂, as suggested by the different volatile mixtures in the Raman analysis sites in our experiments (Fig. 3). The sulfur concentration at sulfide saturation (SCSS) for the magmatic starting materials (anhydrous and mildly hydrated basalt at 1200°C and 150 MPa) was calculated at ca. 0.1 wt % (cf. Fortin et al., 2015). These low solubilities indicate that the mafic host melt is ineffective as a carrier of sulfur at equilibrium conditions. Since only a minor amount of sulfur is soluble in the host melt, the excess sulfur would form a dense, immiscible sulfide 'xenomelt' (Lesher, 2017, 2019). Some of the modified glasses in our experiments contain sulfur exceeding the SCSS (Fig. 2a), which is not surprising since the experiments clearly did not attain equilibrium. We

therefore propose that crustally contaminated melts generated during incipient magma-shale interaction can temporarily carry excess sulfur, which may manifest as a fine suspension of sulfide droplets (i.e. sulfide xenomelts; Fig. 2c and Fig. 4). For the most part, the experiments are visualized here in two dimensions using SEM imaging of the surface. However, inspection of an experiment with extended focus under reflected light revealed a ca. 100 μ m wide zone around shale replete with minute (<5 μ m) sulfides beneath the sample surface (see Supporting Information). This finding supports the notion that the compositional boundary layers in our experiments are a mixture of crustally contaminated silicate melt and fine disseminations of sulfide xenomelts. Our experiments therefore shed light on the origin of sulfide xenomelts in LIP systems that intersect sulfur-bearing sedimentary rocks. Moreover, the experiments support the idea that sulfur-charged contaminated melts can form rapidly, especially in fault-guided conduit systems where magmatic and tectonic brecciation would be common (Hayes et al., 2015) and where they could be forced upward as a slug of sulfur-enriched magma. Since sulfide xenomelts can scavenge metals from mafic magmas, they may eventually lead to the formation of economic ore deposits.

Our experiments highlight that magma–shale interaction in the shallow parts of magma plumbing systems beneath LIPs is an effective means to rapidly mobilize carbon volatiles (Fig. 4). Several major LIPs are characterized by pulsed emplacement of sills into sedimentary basins (e.g. Callegaro et al., 2021; Bédard et al., 2021a, 2021b), which would lead to repeated episodes of magma–sediment interaction and pulsed volatile release throughout the basin. These volatiles could then enter the sub-aerial or sub-aqueous environment via faults and/or hydrothermal vent complexes or breccia pipes, the latter of which are known from the Karoo (Svensen et al., 2007) and Siberian LIP (Svensen et al., 2018) and potentially the Barents Sea region of the HALIP too (Polteau et al., 2016).

Climate warming induced by excess carbon outgassing at LIPs would cause thermal stress in ecosystems and potentially trigger ocean anoxic events (OAEs), a proximal killer in some mass extinction scenarios (Bond & Grasby, 2017). The HALIP was a protracted event spanning more than 40 Myr with a major pulse of continental basaltic magmatism between 135–120 Ma and another at ca. 105–90 Ma (Bédard et al., 2021a, 2021b). These pulses overlap with OAE1a at 120 Ma and OAE2 at 95 Ma. OAE1a was accompanied by sea-surface warming of as much as 8°C (Ando et al., 2008), while OAE2 is considered to be one of the most intense OAEs (Naber et al., 2020). Remarkably, the magnitude of sea-surface warming associated with OAE1a parallels that associated with the Paleocene-Eocene thermal maximum and the Jurassic Toarcian OAE, consistent with the idea that all of these environmental crises share a causal mechanism related to massive release of carbon to the atmosphere-ocean system (Ando et al., 2008).

The driving force for OAE1a and OAE2 is suspected to be volcanic CO₂ outgassing, but the source(s) remain uncertain. Possibilities include dissociation of methane clathrates or volcanism from the Ontong-Java LIP, the Caribbean LIP or the HALIP (Méhay et al., 2009; Midtkandal et al., 2016; Naber et al., 2020). The Barents Sea sill complex, part of the wider HALIP, has been speculated as a source for potentially 20 000 Gt of thermogenic carbon (equivalent to 175 trillion oil barrels) via thermal metamorphism of up to 400 000 km³ of organic-rich sediments due to repeated sill injection (Polteau et al., 2016). This scenario assumes that a maximum of 2 wt % of the carbon in thermal aureoles around sills was expelled (Polteau et al., 2016), which would equate to ca. 4 wt % original TOC assuming that no more than 50% original kerogen is converted to hydrocarbon. Employing a more conservative figure of 1 wt % carbon discharged (=2 wt % original TOC), we arrive at 10 000 Gt thermogenic carbon released from the Barents Sea volcanic basin. Important to note is that these calculations do not consider inorganic carbon that could be effectively mobilized by heating of carbonates. Unfortunately, estimates of carbon release by HALIP intrusions are fraught with uncertainties regarding quantification of parameters such as thermal aureole thicknesses and stratigraphic variability in TOC. However, there are striking parallels between the estimated thermogenic carbon release for the Barents Sea (Polteau et al., 2016) and for the Karoo LIP (Heimdal et al., 2021).

In this study, we employed on-land HALIP sills as a case study and provided experimental constraints that help clarify the mechanisms and rates of magma–shale interaction (summarized in Fig. 4). If we consider devolatilization of a cube of our starting material shale with 500 m side length, ca. 16 Mt of C could be produced. Given the short timescale of shale devolatilization observed here, and that magma volumes of LIPs are on the order of 10⁵ to 10⁷ km³ (Black et al., 2021), it is wholly conceivable that several thousand Gt of C could be generated through repeated episodes of magma-shale interaction proximal to sills in a shallow magma plumbing system. This would moreover be in addition to decarbonation in the distal, lower-temperature parts of metamorphic aureoles (processes that are not directly addressed by our experiments). We therefore suggest that carbon release by magma-sediment interaction at various scales during emplacement of the Canadian Arctic portion of the HALIP could have been similar to estimates for thermogenic carbon release from the Barents Sea (up to 20 000 Gt C; Polteau et al., 2016) or the Karoo (ca. 20 500 Gt C; Heimdal et al., 2021), making the HALIP a strong contender as the causal mechanism for OAE1a and OAE2. However, detailed thermal modelling of the impact of widespread HALIP sill intrusion on Sverdrup Basin sediments would be required to test this hypothesis. A surprising corollary of our work is that entrapment of sulfur in sulfide xenomelts might act to lessen the amount of sulfur released to the atmosphere, which in turn could boost the warming effects of carbon by producing fewer climatecooling sulfurous aerosols. This may account for the general observation (cf. Bond & Grasby, 2017) of LIPs being more strongly associated with climate warming, rather than cooling.

DATA AVAILABILITY

The data underlying this article are available in the article and in its online supplementary material (Supporting Information PDF file and Supporting Data File in Excel format).

SUPPLEMENTARY DATA

Supplementary data are available at Journal of Petrology online.

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